

Obstruction Flat Rigidity of the
CR 3-Sphere

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* Joint with P. Ebenfelt

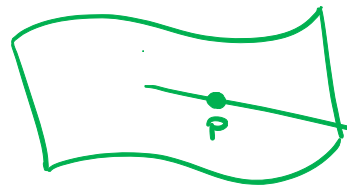
Complex Analysis, Geometry
and Dynamics

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CR 3-Manifolds

$M^3 \subseteq \mathbb{C}^2$ real hypersurface

\leadsto CR structure (M, H, J)



(H_p, \bar{J}_p)
max. compl.
subspace

- (generically) H contact distribution

$\hookrightarrow H = \ker \theta, \quad \theta \text{ 1-form}$
 $\theta \wedge d\theta \neq 0$

- $J: H \rightarrow H, \quad J^2 = -\text{Id}.$

Notation: $(M, H, J) \leftrightarrow (M, T^{1,0}) \leftrightarrow (M, \bar{\partial}_b)$

$$\mathbb{C}H = \underset{i}{T^{1,0}} \oplus \underset{-i}{T^{0,1}} \quad \text{eigenspaces of } J$$

Often $T^{1,0} = \text{span} \{ z_1 \}$

" $\bar{\partial}_b = z_{\bar{1}} = \bar{z}_1$ "



$[T^{1,0}, T^{1,0}] \subseteq T^{1,0}$ trivially.

CR Structure of Unit $S^3 \subseteq \mathbb{C}^2$

$$u = 1 - |z|^2 - |w|^2, \quad \Theta = i\partial\bar{u}|_{TS^3}$$

On S^3

- $\Theta = i(zd\bar{z} + wd\bar{w})|_{TS^3}$ (real)

- $d\Theta = i(dz \wedge d\bar{z} + dw \wedge d\bar{w})$

- $Z_1 = \bar{w}\partial_z - \bar{z}\partial_w$

- $Z_{\bar{1}} = w\partial_{\bar{z}} - z\partial_{\bar{w}}$

- $T = i(z\partial_z + w\partial_w) - i(\bar{z}\partial_{\bar{z}} + \bar{w}\partial_{\bar{w}})$
(Reeb v.f. of Θ)

$$[Z_1, Z_{\bar{1}}] = -iT$$

$$[T, Z_1] = -2iZ_1$$

$$[T, Z_{\bar{1}}] = 2iZ_{\bar{1}}$$

Abstract Deformations

- Gray's theorem: might as well fix H . \rightarrow only vary J .

$$\leadsto \hat{Z}_1 = Z_1 + \underline{\varphi_1}^{\bar{1}} Z_{\bar{1}}$$

$$\varphi_1^{\bar{1}} \in C^\infty(S^3, \mathbb{C})$$

$|\varphi|^2 < 1$
(pointwise)

CR Obstruction Flatness ($\Theta \equiv 0$)

Chern-Moser
vs
Bergman/
Szegö

Fel'derman ('76, '79) proposed the development of CR invariant theory through the study of the formal asymptotics of the following (biholomorphically invariant) Dirichlet problem:

$$\left\{ \begin{array}{l} \bar{\mathcal{T}}(u) = (-1)^n \det \begin{pmatrix} u & u_{\bar{z}^k} \\ u_{z^j} & u_{z^j \bar{z}^k} \end{pmatrix} = 1 \quad \text{in } \Omega \subseteq \mathbb{C}^n \\ \text{bdd., str. } \Psi\text{-conv.} \\ \\ (u > 0 \text{ in } \Omega) \quad \quad \quad u = 0 \text{ on } \partial\Omega \end{array} \right.$$

$v = -\log u$ is a Kähler potential
for ! complete K-E metric

$$(\det(v_{z^j \bar{z}^k}) = e^{(n+1)v})$$

• Cheng-Yau: $\exists!$ solution $u \in C^\infty(\Omega) \cap C^{n+1}(\bar{\Omega})$

• Lee-Melrose: $u \sim \rho + \underline{\Theta} \rho^{n+2} \log \rho + \dots$
a local CR invariant

ρ a Fel'derman
def. function
 $\bar{\mathcal{T}}(\rho) = 1 + O(\rho^{n+1})$

• Graham: $\Theta \equiv 0 \Rightarrow u \in C^\infty(\bar{\Omega})$

\uparrow "CR obstruction density" (Many roles!)

*Physics:
Baeh flat
(conf. gravity)
&
AdS/CFT

Solutions to $\Theta \equiv 0$?


Cauchy-Kowalewski:

Set $\Theta \equiv 0$ up as a Cauchy problem with appropriate initial data.

* Local:

- Graham '87: \exists infinitely many (inequivalent) local real analytic hypersurfaces in \mathbb{C}^n ($\forall n \geq 2$) with $\Theta \equiv 0$.

* Global (compact):

- $u = 1 - \|z\|^2$ on the ball, so the CR spheres are obstruction flat (as are locally spherical structures).
- That is all we know... 

* Rigidity Results:

(non-perturbative)

- C. - Ebenfelt '18: $S^3 \subseteq \mathbb{C}^2$ has no obstruction flat deformations inside \mathbb{C}^2 .

(perturbative)

- Hirachi - et. al. (in preparation): $S^{2n+1} \subseteq \mathbb{C}^{n+1} \dots$ what about outside?

Hope!

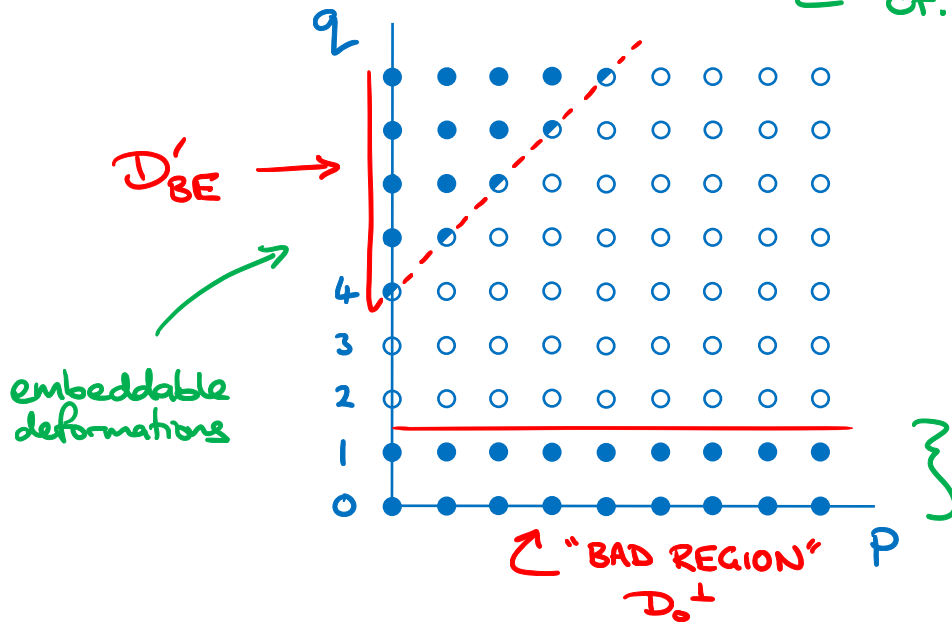
spherical harmonics:

$$H_{p,q} \subseteq P_{p,q}$$

$$(e.g., (z-w)^p (\bar{z}+\bar{w})^q \in H_{p,q})$$

Slice Theorem (C.-Ebenfelt '20):

cf. Burns-Epstein, Bland, Cheng-Lee



Any sufficiently small abstract deformation of the unit $S^3 \subseteq \mathbb{C}^2$ is equivalent to one in this slice (unique mod $\text{Aut}_{\mathbb{C}}(S^3)$).

cf. $\text{PSU}(2,1)$

kernel of the linearized obstruction operator $D\Theta$!

infinite dimensional & complementary to the embeddable deformations!

→ Deformation Theory! Obstructed?

Dashed!

Theorem (C.-Ebenfelt '21):

There is an open neighborhood U of the origin in the slice $\mathcal{D}'_{BE} \oplus \mathcal{D}_0^\perp$ such that the CR structure corresponding to $\varphi, \bar{\tau} \in U$ is obstruction flat if and only if $\varphi, \bar{\tau} = 0$.

* Topology: \mathbb{C}^3 (or better H_{FS}^3)

→ There are no solutions of $\mathcal{O} \equiv 0$ near $(S^3, \text{Holo}, \text{Jnd})$,
embeddable or nonembeddable.

Indication of Proof (by contradiction):

Suppose $\varphi^{(k)} \xrightarrow{H^3_{FS}} 0$ solves $\mathcal{O}(\varphi^{(k)}) = 0 \quad \forall k \in \mathbb{N}$.

① set $\varphi^{(k)} = \varepsilon_k \hat{\varphi}^{(k)}$ with $\|\hat{\varphi}^{(k)}\|_{H^3_{FS}} = 1 \quad (\varepsilon_k \rightarrow 0)$

② $\mathcal{O} = \nabla' \nabla' Q_{11} - iA'' Q_{11}$. Show (WORK) $\|\hat{\varphi}^{(k)}\|_{D'_{BE} H^3_{FS}} \rightarrow 0$
 $\mathcal{O}(\varphi^{(k)}) = 0 \rightarrow \underline{D\mathcal{O}}(\varphi^{(k)})_{D'_{BE}} = \mathcal{F}(\varphi^{(k)}) \quad \uparrow$
6th order in φ
↓
injective on D'_{BE} and inverse gains 6 derivatives in Folland-Stein spaces.
↑ nonlinear terms (involve 6 derivatives but we only control 3!)

$\|\hat{\varphi}^{(k)}\|_{D'_{BE} H^3_{FS}} \rightarrow 0$
 $\|\hat{\varphi}^{(k)}\|_{D_0^\perp H^3_{FS}} \rightarrow 1$

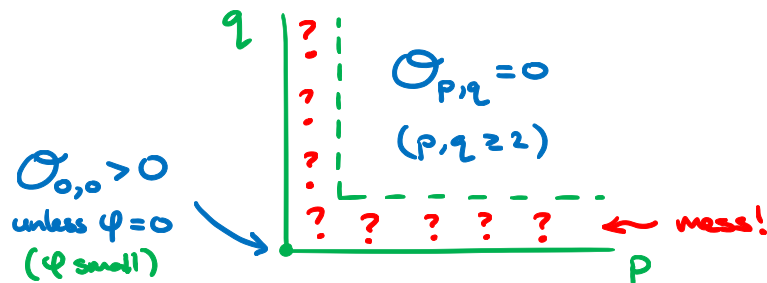
③ $\int_{S^3} \mathcal{O}(\varphi^{(k)}) \sim \varepsilon_k^2 \|\hat{\varphi}^{(k)}\|_{D_0^\perp H^3_{FS}}^2 + \mathcal{O}(\varepsilon_k^{5/2})$

\hookrightarrow so $0 = \lim_{k \rightarrow \infty} \frac{1}{\varepsilon_k^2} \int \mathcal{O}(\varphi^{(k)}) = 1 \quad \nabla \quad \square$

* Formally: we can solve to first order, but not to second order as the $iA'' Q_{11}$ term "wants to be positive" (on average).

Concluding Remarks:

- One can solve $\mathcal{O} \equiv 0 \pmod{\text{coker}(D\mathcal{O})}$,
i.e. $\mathcal{O}_{p,q} = 0$ for $p, q \geq 2$. Every $\dot{\varphi} \in \mathcal{D}_0^\perp$
integrates to a family $\varphi(t)$ of solutions of this
equation (with $\varphi(0) = 0$ & $\frac{d}{dt}|_{t=0} \varphi(t) = \dot{\varphi}$).



- Our result proves linearization instability for the conformal Einstein static universe $S^3 \times \mathbb{R}$ in the space of Bach flat metrics. ("Expected." Cf. Moncrief, Fischer-Marsden, and others in the Einstein case.)

Thank You!

